

# Hyperspectral data quality overview - 2010

This report describes issues that should be considered when further processing any of the 2010 ARSF datasets. The document will be updated over the course of the year, with the latest version available at:

<http://arsf-dan.nerc.ac.uk/trac/wiki/Reports>

## ***Geo-referencing accuracy***

ARSF currently deliver data at level 1 (calibrated sensor data) rather than level 3 (geo-referenced). This allows users to generate level 2 products (e.g. atmospherically corrected radiances) if they wish, and/or map to any projection or datum that suits.

Since ARSF does not currently process data to level 3 except for quality control purposes, instructions for applying precise geometric adjustments to align imagery to other data sets (such as vector overlays) are given to users with delivered datasets. Where a vector overlay or other ground truth information is available, ARSF provide an indication of the average error prior to any such adjustments.

## ***Timing errors***

Correction of the timing offsets between navigation and imagery has been necessary to correct errors in some flight lines. Timing errors cause scan lines to be positioned incorrectly and manifest visually as "wobbles" in the imagery correlated to but out of sync with movements of the aircraft. An example is shown in Figure 1 below.

This is under investigation at various levels and has been raised with Specim.

We endeavour to correct all timing errors prior to delivery. As this is a manual process and relies on finding suitable visible features in the imagery, some errors may still remain. If any are found, please contact us at [arsf-processing@pml.ac.uk](mailto:arsf-processing@pml.ac.uk).



Figure 1a: timing error in an Eagle line



Figure 1b: corrected version of above (0.13 seconds difference)

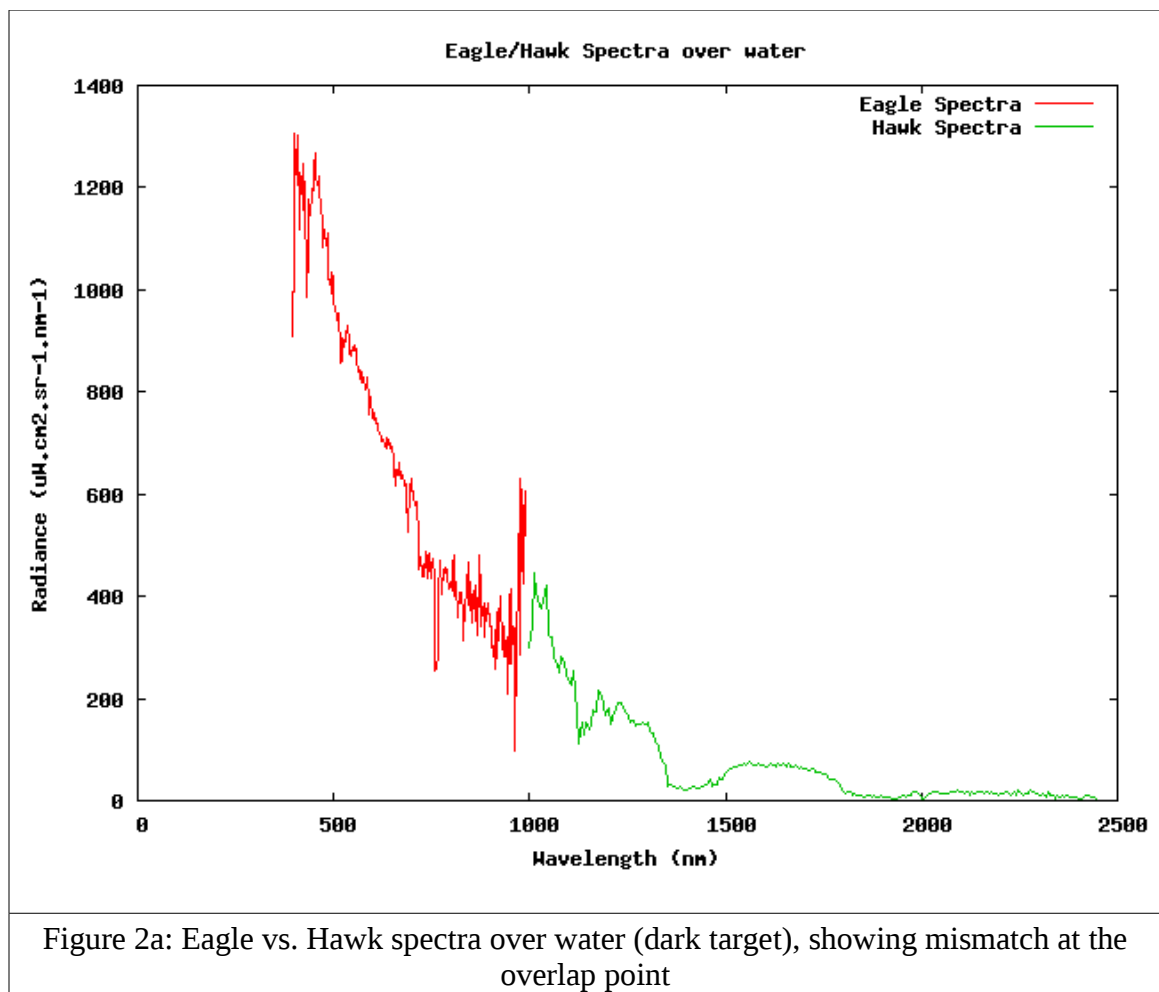
## **Radiometric calibration**

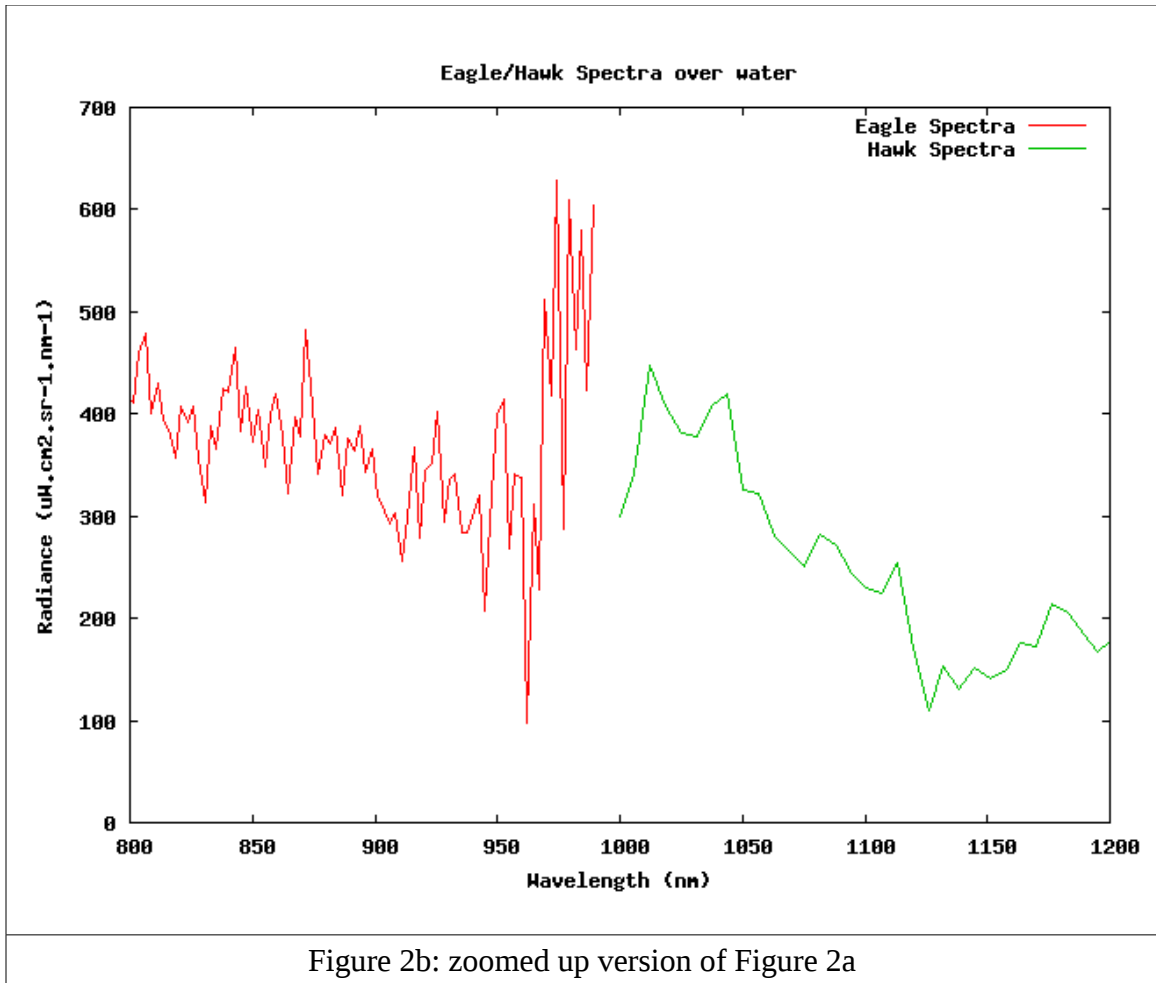
Prior to the start of the 2010 flying season, the Eagle and Hawk instruments were calibrated at Specim. It is also anticipated that the NERC FSF will perform a calibration on the sensors later in the flying season.

## **Quality of 2010 radiometric calibration**

Similarly, as with the 2009 data, the data quality appears to degrade at the low and high wavelength limits of both Eagle and Hawk. For example, a comparison of pixels near the high Eagle wavelengths and low Hawk wavelengths, over dark targets such as water reveals a mismatch (see Figure 2). Caution is advised when examining spectral responses at the edges of the usable range.

Issue tracked at: <http://arsf-dan.nerc.ac.uk/trac/ticket/232>



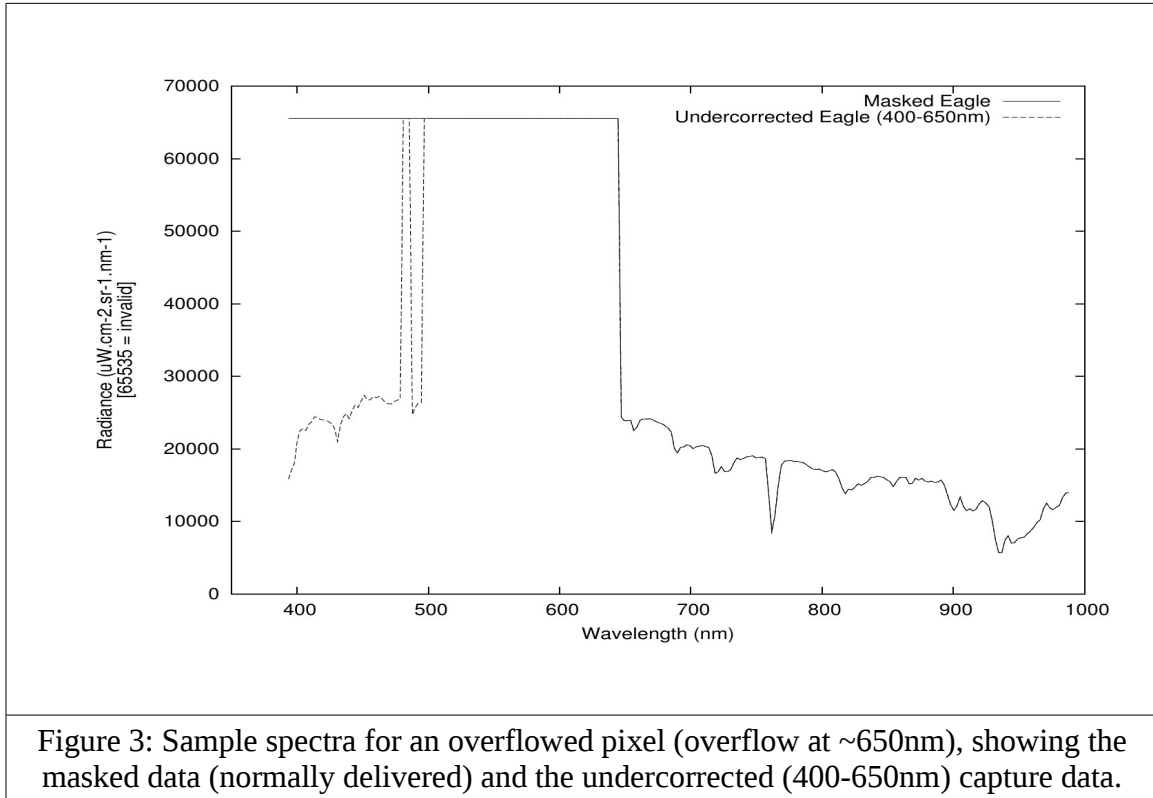


### ***Overflowed pixels***

The instruments have a limited dynamic range and must be set to capture data over the appropriate range of signal strength. For example, if the area of interest is dark, then the instrument will be configured to capture as much low light detail as possible. This configuration is set based on operator experience, the principal investigator's indication of the areas of importance and the prevailing conditions. Inevitably, some pixels are unexpectedly bright – e.g. sunglint over water or part of a cloud. These pixels may exceed the maximum capture level and overflow. Overflowed pixels are marked with an invalid value (65535) in delivered level 1 data. Typically they are not in areas of interest, but should be accounted for.

In Hawk, overflows are marked for just the pixel/band in question. However, Eagle uses a frame transfer CCD, where data are read out in rows. Incoming light continues to accumulate in unread rows during the transfer and is removed by “smear correction” software, which relies on data from one row to correct the next. If a pixel overflows, information is lost and all subsequent pixels in that column cannot be fully corrected. In Eagle, the net effect is that an overflow at 600nm will cause all bluer bands (600nm -> ~400nm) to be under-corrected for that spatial pixel.

When Eagle data with overflows are delivered, we mark as bad all bands following an overflow as they will incorporate some unknown additional light. If your Eagle data contains overflows, the spectrum for that pixel will look like the masked Eagle (solid line) in Figure 3. The dashed line in Figure 3 shows what data were captured, although some are under-corrected – if you wish your data delivered unmasked, please contact us.



## ***Bad pixels***

The Hawk instrument has a number of bad pixels that give inaccurate values. There are different types of error (e.g. constant pixel values, uncorrected offset, duplicating neighbouring pixels, etc), and ~1% of bad pixels are to be expected on the type of CCD used in the Hawk instrument. Specim has provided a list of known bad pixels and these are filled with zero values in delivered data.

While the majority of bad pixels are corrected or blanked out, some have not been detected and may appear in delivered data. These appear in level 1 datasets as straight lines along the direction of flight and as undulating lines in level 3 following the motion of the aircraft (e.g. Figure 4). Typically, they will only affect a single band and are difficult to detect. A complete solution for detecting and removing these will be pursued in 2010.

Issue tracked at: <http://arsf-dan.nerc.ac.uk/trac/ticket/111>

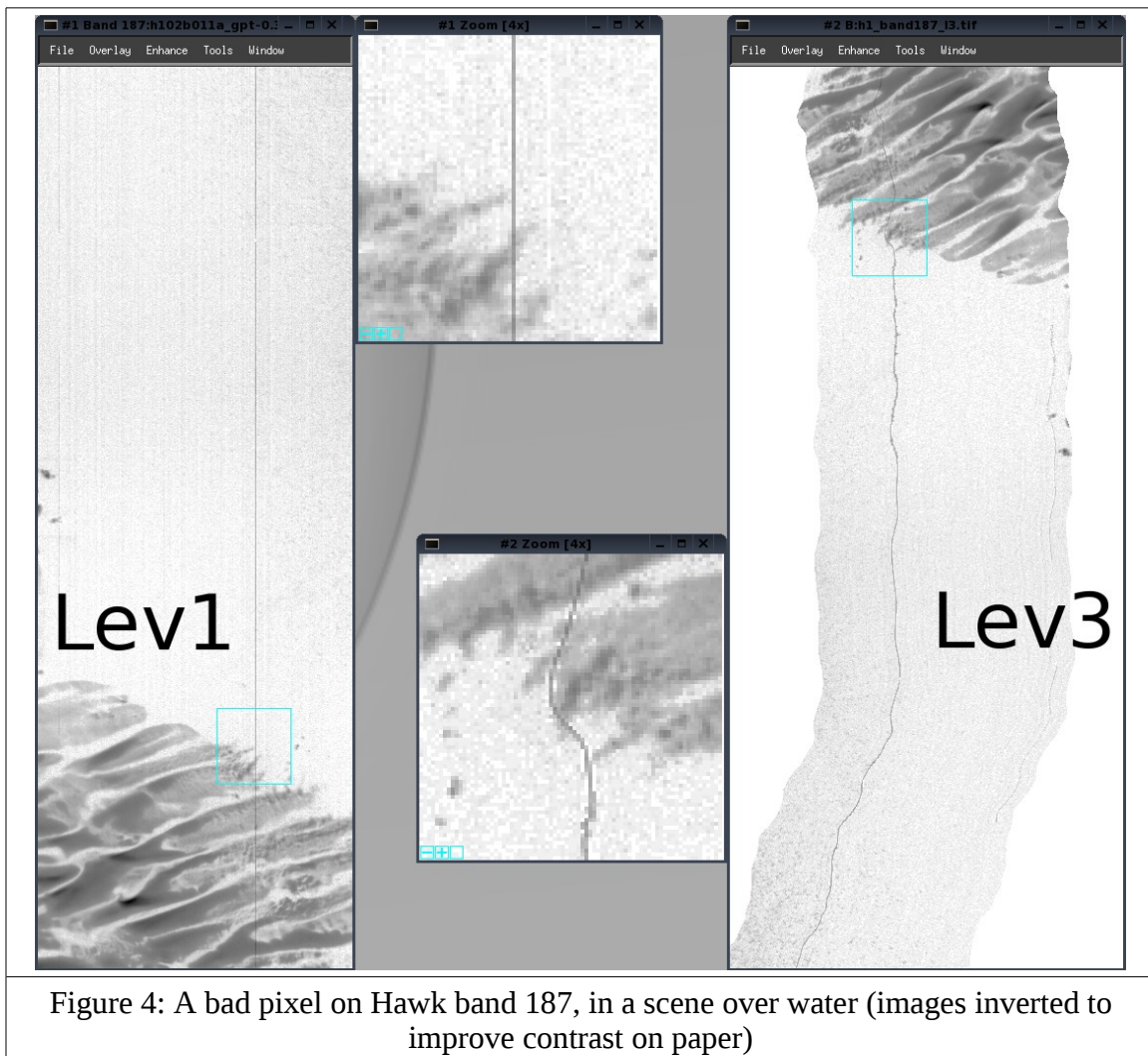


Figure 4: A bad pixel on Hawk band 187, in a scene over water (images inverted to improve contrast on paper)

### ***Eagle dropped frames***

Since late 2009 the Eagle instrument has been suffering from dropping frames at regular intervals. It is currently thought that this is due to an issue with the frame grabber card. This results in regular gaps in the data coverage in the level 1 data (Figure 5). Small gaps in the level 1 data will usually get filled in by interpolation at level 3 stage, though large gaps often remain.

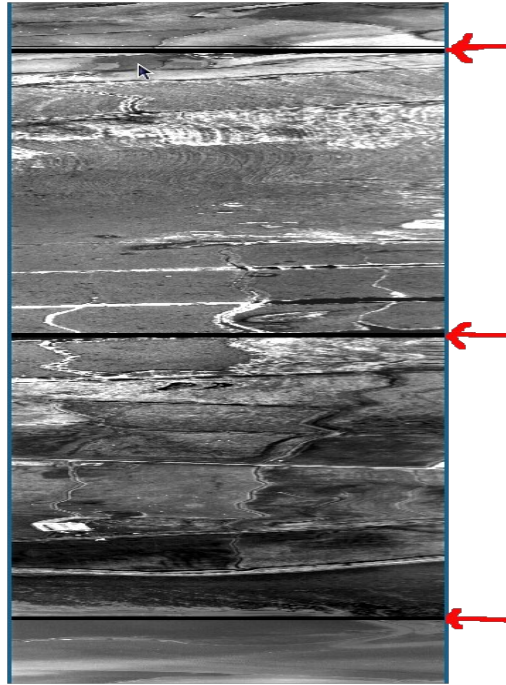


Figure 5: Example Eagle level 1 data with highlighted regular dropped frames